# UNIFIED FIELD THEORIES OF EINSTEIN'S TYPE DEDUCED FROM A VARIATIONAL PRINCIPLE: CONSERVATION LAWS. 

Dedicated to Professor A. Kawaguchi on the occasion of his 70th birthday.

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1. Introduction. In previous papers [4], ${ }^{11}$ [5], [6] we have studied the following generalization of Einstein's unified theory of 1950. The space-time is assumed to be a four dimensional differentiable manifold endowed with a non-symmetric affine connection $\Gamma_{j k}^{i}$ and a non-symmetric covariant tensor $g_{i j}$. The most general covariant tensor $T_{i j}$, which depends only on the connection $\Gamma_{j i}^{i}$ and its first partial derivatives and it is at most of second degree as function of $\Gamma_{j_{k}}^{\prime}$, is the tensor (2.13) where $\alpha, \beta, \ldots$, $\nu$ are arbitrary constants. Then we form the density $T_{i A} g^{i / 4}|g|^{1 / 8}$ and deduce the field equations from the corresponding variational principle. The field equations depend on the set of constants $\alpha, \beta, \ldots, \nu$. The classical Einstein's theory corresponds to $\alpha=1, \beta=\gamma=\cdots=\nu=0$. Particular cases have been considered by M. A. Tonnelat ([7], Note II) and ([8], p. 351-363) where related works of Nguyen Phong Chau (1963), J. Lévy (1959) and L. Bouche (1961) are mentioned. In [6] we have analyzed the conditions of the constants $\alpha, \beta, \ldots, \nu$ for the field equations to be invariant by $\lambda$-transformations or for $T_{i j}$ to be a pseudo-hermitian tensor. In the present paper we shall give some complements of the general theory and, in particular, we establish the conservation laws or conservation identities which are satisfied for any set of variables ( $\Gamma_{j b}^{j}, g_{i j}$ ) which satisfies the field equations.

Though we use small changes in the notations, the main references for the concepts and formulas in the sequel are the books of A. Lichnerowicz [3] and M. A. Tonnelat [7], [8].
2. Notations and field equations. Let $\Gamma_{j k}^{j}$ be an affine connection and let

$$
\begin{equation*}
\Delta_{j k}^{i}=\frac{1}{2}\left(\Gamma_{j k}^{i}+\Gamma_{k j}^{\prime}\right), \quad S_{j k}^{i}=\frac{1}{2}\left(\Gamma_{j k}^{i}-\Gamma_{k j}^{i}\right) \tag{2.1}
\end{equation*}
$$

be its symmetric and skewsymmetric parts ( $\mathcal{S}_{j k}^{\prime}$ is a connection and $S_{j k}^{j}$ is the tensor of torsion). Following Einstein we set

$$
\begin{equation*}
S_{i}=S_{i k}^{\dagger} . \tag{2.2}
\end{equation*}
$$

Let $g_{i j}$ be a non-symmetric tensor. If $g$ denotes the determinant of $g_{i j}$, assumed $\neq 0$, we introduce the densities

$$
\begin{equation*}
\mathscr{G}_{i j}=g_{i j}|g|^{1 / 8}, \quad \oiint_{i j}=\frac{1}{2}\left(\mathbb{G}_{i j}+\mathscr{F}_{j i}\right), \quad \mathfrak{J}_{i j}=\frac{1}{2}\left(\mathbb{G}_{i j}-\mathscr{G}_{j i}\right) . \tag{2.3}
\end{equation*}
$$

For any tensor or density of second rank, Einstein [1] introduced the mixed

1) Numbers in brackets refer to the references at the end of the paper.
covariant derivative，obtained when one differentiates the first index with respect to $\Gamma_{j h}^{j}$ and the second index with respect to $\Gamma_{j h}^{j}=\Gamma_{h j}^{i}$ ．We will denote this mixed covariant derivative by a vertical bar．For instance we put
where a comma denotes ordinary partial derivative．Denoting by a semi－colon the ordinary covariant derivative with respect to the connection $\Gamma_{j k}^{j}$ ，we have

Notice that for symmetric densities $\Phi^{\boldsymbol{1}}=\Phi^{\boldsymbol{1}}$ we have
and for skewsymmetric densities $\mathfrak{F}^{\mathbf{4}}=-\mathfrak{F}^{\boldsymbol{4}}$ ，

In particular，we have

$$
\begin{align*}
& \Phi^{i n}{ }_{16}=⿹^{n i}{ }_{16}+2 S_{m} ⿹^{m A}  \tag{2.8}\\
& \mathfrak{F}^{A}{ }_{k}=-\mathfrak{F}^{A}{ }_{k}+2 S_{i m}^{A} \mathfrak{g}^{i m}+2 S_{m} \mathfrak{f}^{m A} . \tag{2.9}
\end{align*}
$$

Notice also the formulas

$$
\begin{align*}
& \Phi^{n i}{ }_{11}=\Phi^{A i}{ }_{6}+\Gamma_{m,}^{A} \Phi^{m i}-S_{1} \phi^{A i} . \tag{2.10}
\end{align*}
$$

Though in general we shall use densities instead of tensors，we state the follow－ ing formula：
which holds good for any covariant derivative．By means of（2．12）the field equa－ tions in the sequel may be expressed in terms of the tensor $g_{i j}$ ．

The most general covariant tensor $T_{i a}$ ，which depends only on the connection $\Gamma_{\text {in }}^{d}$ and its first partial derivatives and it is at most of second degree as functions of $\Gamma{ }_{3}^{\prime}$, ，is

$$
\begin{align*}
& +\omega_{i ; k}+\phi S_{k ; ~}+\mu S_{m} S_{i k}+\nu S_{i} S_{\lambda} . \tag{2.13}
\end{align*}
$$

where $R_{A}$ is the Ricci tensor

For a proof， 800 ［6］．
The variational principle

$$
\begin{equation*}
\delta \int T_{a n} \sigma^{d a} d \tau=0 \tag{2.15}
\end{equation*}
$$

（where $d \tau=d x_{1} \wedge d x_{1} \wedge d x_{1} \wedge d x_{1}$ ）gives rise to the field equations
（sea rol．whore equations（19）are cloarly misprinted），where

Unified field theories of Einstein's type deduced from a variational principle. 385
(2.17)

By means of (2.12) it is easy to express these equations in terms of the tensor $g^{00}$ instead of the density Gr". For instance, for the classical case $\alpha=1, \beta=\gamma=$ $\cdots=\nu=0$, we get
where $\gamma_{i}=\left(|g|^{1 / 2}\right), 6$ in accordance with Lichnerowicz [3].
The expression (2.17) takes a simple form if the constants $\alpha, \beta, \gamma, \ldots, \nu$ are such that
(2. 18) $\quad \alpha \neq 0, \quad \alpha+\gamma \neq 0, \delta=0, \quad(2 \phi+\mu) \alpha+(c+\phi) r=0$,
a set of conditions which we will assume satisfied from now on.
In this case it is useful to introduce the new connection

$$
\begin{equation*}
L l_{r}=\Gamma_{i_{r}}+\frac{1}{3}(2-(c+\phi) / \alpha) d\left\{^{2} S_{r}-\frac{1}{f}((c+\phi) / \alpha) \delta_{r}^{2} S_{i},\right. \tag{2.19}
\end{equation*}
$$

which is such that

$$
\begin{equation*}
L_{i}=\frac{1}{2}\left(L_{i f}^{f}-L_{i 4}^{i}\right)=0 \tag{2.20}
\end{equation*}
$$

Then the first equations (2.16) split into (see [6])

$$
\begin{align*}
& +\frac{1}{8}(-2 \alpha+2 \beta-\gamma) \delta^{\circ} \mathfrak{z}^{\prime \prime}{ }^{\prime} \text {, }=0 \text {, } \tag{2.21}
\end{align*}
$$

$$
\begin{align*}
& -\{(c+\phi)(2 \alpha-\varepsilon-\phi) / \alpha+\delta+3 \nu\} S_{i} \Phi^{\alpha 1}=0, \tag{2.22}
\end{align*}
$$

where the mixed covariant derivatives refer to the connection $L!$.
From (2.21) having (2.8) and (2.9) into account, we deduce

Hence, according to (2.10) and (2.20), the equations $a_{i}^{i f}-a_{i}^{i}=0$ are identically aatisfied and that justifies the addition of (2.22) as field equations. On the other side (2.21) gives

Hence, from (2.23) and (2.24), asauming $\alpha+\gamma \neq 0$, we have

$$
\begin{equation*}
\mathfrak{F}^{\alpha d}(L)=L_{m} \delta^{\prime m}+\mathfrak{F}^{a d}, \tag{2.25}
\end{equation*}
$$

Substituting (2.25) and (2.26) into (2.22) and having (2.18) into account, we get the interesting relation

$$
\begin{equation*}
A \mathfrak{Z}^{a i}, i+B S_{i} \mathscr{Q}^{i i}=0 \tag{2.27}
\end{equation*}
$$

where

$$
\begin{align*}
& A=2 \alpha-5 \beta+5 \beta(\varepsilon+\phi) / \alpha-s-4 \phi+\gamma,  \tag{2.28}\\
& B=-(\varepsilon+\phi)(2 \alpha-s-\phi) / \alpha-\delta-3 \nu . \tag{2.29}
\end{align*}
$$

Notice that by means of the connection $L_{i r}$ the tensor $T_{i n}$ writes

$$
\begin{equation*}
T_{i k}=T_{i h}(L)+\frac{1}{3} A\left(S_{i, A}-S_{h, i}\right)-\frac{1}{3} B S_{i} S_{A}, \tag{2.30}
\end{equation*}
$$

and the second set of field equations (2.16) writes

$$
\begin{equation*}
T_{i n}(L)+\frac{1}{3} A\left(S_{i, \lambda}-S_{h, i}\right)-\frac{1}{3} B S_{i} S_{k}=0 . \tag{2.31}
\end{equation*}
$$

The field equations are (2.21), (2.22) and (2.31) from which the relations (2.25), (2.26) and (2.27) follow. The unknowns are $L_{i r}, \operatorname{G}^{i d}, S_{i}$. The components $\Gamma_{i r}$ are then given by (2.19) from which the relation $\Gamma_{i q}-\Gamma_{i}^{i}=2 S_{i}$ follows.
3. Particular cases of the fleld equations. Taking account of (2.27) we see that there are two important particular cases to be considered:
a) $B=0, A \neq 0$. Then (2.27) gives

$$
\begin{equation*}
\mathfrak{z}^{a}=0, \tag{3.1}
\end{equation*}
$$

and equations (2.21) give

$$
\begin{equation*}
\alpha \mathbb{G}{ }_{1 r}(L)+\gamma \mathfrak{F}^{q \cdot}{ }_{1 r}(L)=0 . \tag{3.2}
\end{equation*}
$$

Equations (2.31), if $B=0$, may be written as

$$
\begin{equation*}
T_{(i \alpha)}(L)=0, \quad T_{[\{A], j}(L)+T_{[A j], k}(L)+T_{[j i], k}(\dot{L})=0, \tag{3.3}
\end{equation*}
$$

where () denotes the symmetric part and [] the skewsymmetric part of the tensor $T_{i k}(L)$. The field equations are (3.1), (3.2) and (3.3). When $\gamma=0$, this system reduces to the so called "weak system" of Einstein. Thus, we have proved that any tensor $T_{i n}(2.13)$ such that the constants $\alpha_{1} \beta, \gamma, \ldots, \nu$ satisfy the conditions (2.18) and $B=0, \gamma=0$ ( $B$ given by (2.29)) gives rise to the field equations of the weak system of Einstein.
b) $A=0, B \neq 0$. Assuming that the determinant $\left\|\mathscr{Q}^{d i}\right\| \neq 0,(2.27)$ gives

$$
\begin{equation*}
S_{i}=0 \tag{3.4}
\end{equation*}
$$

Then, according to (2.19) we have $L \ell_{r}=\Gamma_{\text {r. }}^{\prime}$. Having into account (2.18), (2.25), (2.26) and (3.4), it follows that equations (2.22) are identically satisfiod, so that the field equations reduce to (2.21), (2.31) and (3.4) which may be written as

$$
\begin{align*}
& S_{i}=0, \quad T_{i \mathrm{~h}}=0 . \tag{3.5}
\end{align*}
$$

The more simple case corresponds to $\gamma=0, \beta=0$. Then the system (3.5) takes the simple form
where

Unified field theories of Einstein's type deduced from a variational principle.

$$
T_{i n}=\alpha R_{i n}+(2 \alpha-4 \phi) S_{i ; h}+\phi S_{n ; i}-2 \phi S_{m} S_{i k}^{m}+\nu S_{i} S_{n} .
$$

These field equations are valid for any set of constants $\alpha \neq 0, \phi, \nu$. We have applied that in the present case we have $\varepsilon=2 \alpha-4 \phi, \mu=-2 \phi$. For instance, the particular tensors ${ }^{1} R_{i h},{ }^{2} R_{i n}$ considered by Tonnelat ([7], pp. 129-130) belong to this class. Indeed, ${ }^{1} R_{i n}$ corresponds to $\alpha=1, \phi=\frac{2}{3}, \nu=0$ and ${ }^{2} R_{i n}$ corresponds to $\alpha=1, \phi=\frac{1}{3}, \nu=-\frac{1}{3}$.

The tensor ${ }^{3} R_{i n}$ of Tonnelat ([7], pp. 129-130) corresponds to $\alpha=1, \beta=\frac{1}{2}$, $\gamma=0, \delta=0, c=\frac{2}{3}, \phi=\frac{1}{3}, \mu=-\frac{2}{3}, \nu=-\frac{1}{3}$ and hence $A=0, B=0$. Equations (2.27) are identically satisfied and the field equations reduce to (2.21) and (2.31) which in this case may be written as
c) As a last example we consider the Einstein tensor (see [1])

$$
\begin{equation*}
E_{i h}=-\frac{1}{2}\left(\Delta_{i m, h}^{m}+\Delta_{h m, i}^{m}\right)+\Gamma_{i h, m}^{m}+\Gamma_{i k}^{i} U_{i m}^{m}-\Gamma_{i,}^{m} \Gamma_{m h}^{i}, \tag{3.8}
\end{equation*}
$$

which corresponds to

$$
\begin{equation*}
\alpha=1, \quad \beta=\frac{1}{2}, \quad t=1, \gamma=\delta=\phi=\mu=\nu=0 . \tag{3.9}
\end{equation*}
$$

We have $A=1, B=-1$. The connection (2.19) writes

$$
\begin{equation*}
L_{i r}^{i}=\Gamma_{i r}^{q}+\frac{1}{3} \delta\left\{S_{r}-\frac{1}{3} \delta_{r}^{q} S_{i},\right. \tag{3.10}
\end{equation*}
$$

and the field equations (2.21) write

$$
\begin{equation*}
- \text { Gra }{ }_{1 r}(L)+\frac{1}{3} \delta_{r}^{2} \mathfrak{q ^ { a s }}, 4-\frac{1}{3} \delta_{r}^{4} \mathfrak{V ^ { a s }}, 4=0 . \tag{3.11}
\end{equation*}
$$

The field equations (2.22) are equivalent to (2.27), i.e.

$$
\begin{equation*}
\mathfrak{X}^{a i}{ }_{i}=S_{i} \mathscr{母}^{a d} \text {. } \tag{3.12}
\end{equation*}
$$

The field equations are thus (3.11), (3.12) and $E_{i n}=0$. Adding the condition $S_{i}=0$ we get the "strong system" of Einstein grr" ${ }_{1 r}=0, S_{i}=0, E_{i \mathrm{~h}}=0$ which, however, is not deducible from a variational principle.
4. Conservation laws. To get the identities of conservation we will follow a similar approach to that of Lichnerowicz and Weyl for the case $\alpha=1, \beta=\gamma=\ldots$ $=\nu=0$ (soe Lichnerowicz [2]).

Let $C$ be a domain of the space-time of boundary $\partial C$ and let $\xi^{\prime}$ be a vector field which vanishes on $\partial C$. Consider

$$
\begin{equation*}
I=\int_{C}{ }^{i \operatorname{bih}} T_{i n} d \tau=\int_{C} \mathfrak{x} d \tau \tag{4.1}
\end{equation*}
$$


The Lie derivative with respect to the field $\xi^{4}$ is

$$
\begin{equation*}
L_{l} I=\int_{C} L_{l} \mathfrak{N} d \tau=\int_{0}\left(\mathfrak{x} \xi^{\prime}\right) \cdot d \tau \tag{4.2}
\end{equation*}
$$

By means of the Stokes' theorem this intogral transforms into an integral exrended on $\partial C$, which is zero since $\xi^{4}$ vanishes on $\partial C$. Thus we have

$$
\begin{equation*}
L_{i} I=0 \tag{4.3}
\end{equation*}
$$

On the other side, recall that the Lie derivative of $\Gamma_{j k}^{j}$ is a mixed tensor of contravariant valence 1 and covariant valence 2 , given by

$$
\begin{equation*}
L_{\ell} \Gamma_{j k}^{i}=\xi_{, j k}^{i}+\xi^{m} \Gamma_{j k, m}^{i}-\xi^{i}{ }_{, m} \Gamma_{j k}^{m}+\xi^{m},{ }_{, j} \Gamma_{m k}^{i}+\xi^{m}{ }_{, k} \Gamma_{j m}^{i} \tag{4.4}
\end{equation*}
$$

(see Yano [9]). Thus, assuming that the vector field $\xi^{i}$ and its first and socond derivatives vanish on $\partial C$, we have, on $\partial C, L_{8} \Gamma_{j h}^{j}=0$. According to the variational principle from which the field equations are deduced, this condition implies that

$$
\int_{c}\left(L_{l} T_{i \alpha}\right) \not \overleftarrow{S}^{i \lambda} d \tau=0,
$$

and (4.3) gives

$$
\begin{equation*}
L_{l} I=\int_{C} T_{i \Lambda} L_{f}^{\left(\sigma^{i \alpha}\right.} d \tau=0 \tag{4.5}
\end{equation*}
$$

As is well known (see Yano [9]) we have
and hence

Substituting this expression in (4.5), applying then the Stokes' theorem and having into account that the vector field $\xi^{i}$ is an arbitrary vector field which vanishes on $\partial C$ (together with its first and second derivatives), we get

$$
\begin{equation*}
\left(T_{\Delta s} G^{m \lambda}+T_{i s} G^{(J m}\right)_{\cdot m}-G^{i n} T_{i h, 4}=0 \tag{4.8}
\end{equation*}
$$

Putting
(4.8) may be written as

$$
\begin{equation*}
\mathbb{B}_{,}^{m}, \mathrm{~m}+\frac{1}{2} T_{1 k} \mathbb{G}^{4 n}{ }_{, 0}=0, \tag{4.10}
\end{equation*}
$$

which is the first form of the four identities of conservation.
These identities refor to the connection $\Gamma \mathbf{l}_{r}$. If we want to introduce the connection Lir (2.19) which gives rise to the field equations (2.21) and (2.22), notice that subetituting the expression (2.30) of $T_{i k}$ into (4.9) we get

$$
\begin{align*}
& \left.-\frac{1}{j} B S_{j} S_{k}\right) G^{k m}-\frac{1}{8} \delta_{j}^{m i d}{ }^{1 j}\left(T_{i j}(L)+\frac{1}{j} A\left(S_{i, j}-S_{j, i}\right)-\frac{1}{3} B S_{i} S_{j}\right) \tag{4.11}
\end{align*}
$$

$$
\begin{aligned}
& \left.-\frac{1 \delta_{i}^{m}}{} A\left(S_{i, j}-S_{j, i}\right)\right)^{4 j}-\frac{1}{3} B S_{,} S_{k} \Phi^{m \hbar}+1 B \delta_{j}^{m} \Phi^{4} S_{i} S_{j} .
\end{aligned}
$$

Putting

and having into account the value of $188^{\prime \prime}$, an easy calculation shows that (4.10) may be written as

Unified field theories of Einstein's type deduced from a variational principle.
(4. 13) $\mathfrak{R}_{0, m}^{m}+\frac{1}{2} T_{i j}(L) \mathscr{G}^{i j}{ }_{, 0}+\frac{1}{3} S_{h, 0}\left(A \mathcal{F}^{n k}{ }_{, 4}+B S_{i} \mathscr{Q}^{h i}\right)-\frac{1}{8} A S_{i, h \mathcal{Y}^{h i}}{ }_{, 4}=0$, which is the second form of the identities of conservation.

Notice the relation


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